A New Radioactive Decay Code for Geant4



S. Hauf, M. Kuster, M. Batič, Z. W. Bell, D. H. H. Hoffmann, P. M. Lang, M. G. Pia, G. Weidenspointner, A. Zoglauer

class G4RadioactiveDecay : public G4VRestDiscreteProcess { PU/OGLAICCollectiveDecay(const G4String& processName="RadioactiveDecay"); statisticalApproach=tnue;RadG4booly", G4bool taccountForRecoil=true, G4bool statisticalApproach=true, G4bool tdecayChainAnalysis=false); G4VParticleChange * VecOt(Loof G4TroSepters tradgeterl Stra) 1 (GetVREmission()_{G4i}{returnsvrDoEmission;} G4bool GetVREmission(G4int Z, G4int A) {return vrDoEmission[Z][A];} std::unordered_map<G4int, std::unordered_map<G4int, G4bool> > GetVREmission() {return vrDoEmission;} void SetVRN0ForNucleus(G4int Z, G4int A, G4double N){N0[2][A]=1.; vr=true; vrAll=true; ntimesVR=N;}

Contents



Radioactive decays in Monte-Carlo Simulations
 Validation and Verification

Decay Chaine & Activation

Decay-Chains & Activation



- Radioactive decays play a role in diverse fields: material sciences, astrophysics, nuclear physics, biophysics, nuclear proliferation monitoring, homeland security...
- Usually: observable and/or background source
- Problem: as experiments get more complex radiation characteristics and background are not easily predictable
- Common solution: Monte Carlo simulations



 Radioactive decay code implementations exist in popular general purpose Monte-Carlo frame-works:

MCNP

 Delayed gammas from tabulated data
 Possibility to integrate specialized tools such as ORIGEN2 or CINDER90, adding activation, processing of decay chains
 Alpha and beta-particles from tabulated input
 Summary statistics

FLUKA

Generates and transports gamma and beta radiation
Alpha decays only added in newest version
Decay chains and activation
NNDC library based on ENSDF
Summary statistics

Geant4

- ► Generation of alpha, beta and gamma radiation
- ►Three distinct processes involved
- Decay library based on ENSDF
- Decay chains
- Activation present as distinct process
- Per-decay simulation



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Per-decay simulation

Experimental requirements (per-decay, summary) defines framework



- Radioactive decay code implementations exist in popular general purpose Monte-Carlo frame-works
- Experimental requirements (per-decay, summary) defines framework
- •What if other requirements favor other frame-works?





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- Radioactive decay code implementations exist in popular general purpose Monte-Carlo frame-works
- Experimental requirements (per-decay, summary) defines framework
- •What if other requirements favor other frame-works?
- Solution: an implementation which allows for physically accurate per-decay sampling AND statistical sampling when the outcome of many decays is of interest.



Software Requirements for a **Radioactive Decay Simulation**



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Decay an unstable nucleus

- Sample discrete decay emission: alpha, neutrinos (EC)
- Sample continuous decay emission: Beta-Fermi-function (e, neutrinos)
- Generate daughter nucleus
- Take care of kinematics: energy- and momentum conservation

Deexcite daughter nucleus

- Sample gamma, X-rays and conversion electrons resulting from nuclear deexcitation
- Sample fluorescence and Auger-electrons due to shell vacancies
- Handle Decay chains efficiently
- Handle long term activation

The per-Decay approach of the *current* Geant4 Radioactive Decay Code



Original implementation: F. Lei & P. Truscott (Qinetic), Geant4.9.4-p04 **Decay an unstable nucleus** Functional per-decay approach with a few ✓ Sample discrete decay emission: alpha, neutrinos (EC) inconsistencies and missing features Based on popular ENSDF library \checkmark Sample continuous decay emission: Beta-Fermi-function (e, neutrinos) Adapt what is good ✓ Generate daughter nucleus ✓ Take care of kinematics: energy- and momentum conservation **Deexcite daughter nucleus** ✓ Sample gamma, X-rays and Improve where necessary conversion electrons resulting from atomic deexcitation Sample fluorescence and Augerelectrons due to shell vacancies ✓ Handle Decay chains efficiently Add what is missing Handle long term activation

A *new* Radioactive Decay Code for Geant4



- ENSDF-based library only data source
- Compatible with NuDAT
- Efficient sampling of many decays

Per-decay approach based on adapted F. Lei & P. Truscott implementation.

- Corrected fluorescence sampling bug (independently corrected in G4.9.5)
- Corrected energy conservation problem (present in G4.9.5)
- Flexible delegation to deexcitation and relaxion processes, i.e. changes can easily be included
- Flexible Beta-Fermi function implementation
- Efficient handling of decay chains sampling only isotopes of interest
- Radioactive background sources
- Provision of bookkeeping classes for long term activation

Decay an unstable nucleus

- ✓ Sample discrete decay emission: alpha, neutrinos (EC)
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- ✓ Generate daughter nucleus
- ✓ Take care of kinematics: energy- and momentum conservation

✓ Deexcite daughter nucleus

- ✓ Sample gamma, X-rays and conversion electrons resulting from atomic deexcitation
- ✓ Sample fluorescence and Augerelectrons due to shell vacancies
- Handle Decay chains efficiently
- Handle long term activation
- New statistical sampling approach



Verification of both Codes with ENSDF

- Both codes uses a library based on the Evaluated Nuclear Structure Data File (ENSDF)
- Verification/Validation data from NNDC supplied program RADLST
- ►~3000 isotopes&levels, broader than any previous verification
- Additionally validation against own data from HPGe detector measurements
 - Experimental data
 - Derived from models

consistent database

single

new implementation





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Intensity Verification Gamma Rays (>30 keV)





Intensity Verification Electrons (<30 keV)





Energy Verification All radiation types (>30 keV)





Computational Performance Single decays





Single decay performance

Efficient treatment of Decay Chains



- Usual procedure: extended Bateman solution with branching and recursive implementation
- New approach: algebraic solution developed by M. Amaku *et al.*
- Advantage: only need to retrieve decay properties once, reuse for any initial isotope number
- More details in next talk on long-term activation



Summary up to now...



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- Radioactive decay in Monte-Carlo simulations follows two flavors:
 - Per-decay simulation
 - Statistical summary of many decays
- Experimental requirements for both approaches exists
- Currently, no general purpose Monte-Carlo framework offers both

Newly developed radioactive decay code allows for both approaches

- Verification shows: it is more accurate with respect to ENSDF library
- Computational performance tests show: it is faster than current Geant4 code concerning single decays and especially decay chains
- Additional bonus: modern code design, i.e. more maintainable
- Additional bonus: useful long term activation implementation (see next talk)
- And: bug fixes concerning energy/momentum conservation, decay emission production
- Additional validation with measurements from HPGe detector (N14-240)

The Influence of In-Orbit Activation Buildup on the of the ATHENA Wide Field Imager



S. Hauf, M. Kuster, D. H. H. Hoffmann, P. M. Lang, S. Neff, M. G. Pia, G. Weidenspointner, A. Stefanescu, L. Strüder



Contents

- A short introduction to ATHENA and the Wide Field Imager
- In-orbit background sources
- The prompt background
- Handling long-term activation in the new Geant4 Decay Code
- Results of the delayed background simulation

A Telescope for High ENergy Astrophysics - ATHENA

Planned next generation X-ray telescope

Designated successor for XMM Newton (max. until 2018) and Chandra (max. until 2015)

High sensitivity

→ max. background: 10⁻³ cts keV⁻¹ cm⁻² s⁻¹

- Good spatial resolution < 10^e
- Good temporal resolution ~ 10 µs
- ▶Good spectral resolution ~ 3eV @ 1keV

►DePFET-based Wide Field Imager imaging spectroscopy at energies between 0.1-15 keV on 640x640 pixel

The ATHENA Background Environment

- Position: 2nd Lagrangian
- Benefit: allows continuous observations not affected by passes through Earth's shadow and radiation belts
- Downside: unprotected by geomagnetic field. Harsh cosmic ray environment, solar protons and X-ray @ energies in the keV-TeV range
- Satellite and detector components will get activated
- What background is to be expected?
- Can't replicate environment on Earth – need Monte-Carlo Simulations

Background-Image: Wikipedia, Orbit-Plot: Bacons et Al, 20

ATHENA Simulation Geometry in Geant4

Validation of the Simulation Environment with XMM Newton EPIC-pn Data

- Simulation of EPIC-pn recorded filter-wheel closed data, i.e. only cosmic rays and secondaries hit sensitive area
- Details only for PCB, everything else was an aluminum box
- Magnitude and shape of background in general agreement except for Ni-line

Comparison simulation - measurement

Validation of the Simulation Environment with XMM Newton EPIC-pn Data

Simulation of EPIC-pn **EPIC-pn quantum efficiency** recorded filter-wheel closed 1.2 data, i.e. only cosmic rays and secondaries hit sensitive area 1.0 Details only for PCB, 0.8 everything else was an Quantum efficiency aluminum box 0.6 1.02 Quantum efficiency Magnitude and shape of 1.0 0.98 background in general 0.4 0.96 agreement except for Ni-line 0.94 0.92 0.2 1.83 1.87 1.82 1.84 1.85 1.86 Energy [keV] 0.0 Dev. [%] -5 -10 10 1 Energy [keV] Simulation model of the EPIC-pn PCB (backside shown)

Lessons learned from the prompt Background of the ATHENA WFI

Back to the new Decay Code: how decay chains and activation are handled

- Usual procedure: extended Bateman solution with branching and recursive implementation
- New approach: algebraic solution developed by M. Amaku et al. (building upon work by R. J. Onega, D. Pressyanov, L. Moral and A. F. Pacheco, T. M. Senkov and D. Yuan and W. Kernan)
- Advantage: only need to retrieve decay properties once, reuse for any initial isotope number
- By defining chain members of interest and taking into account time spans only necessary isotopes are actually decayed

$$\frac{dN_1(t)}{dt} = -k_1 N_1(t)$$

$$\frac{dN_2(t)}{dt} = \lambda_1 N_1(t) - k_2 N_1(t)$$

$$\vdots$$

$$\frac{dN_n(t)}{dt} = \lambda_{n-1} N_{n-1}(t) - k_n N_n(t)$$

$$k_i = \lambda_i + \alpha_i.$$

Decay Activation constant parameter

Isotopes in chain after time *t*

 $\mathbf{N}(t) = e^{\Lambda t} \mathbf{N}(0)$

[1] M. Amaku, P. R. Pascholati, and V. R. Vanin, "Decay chain differential equations: Solution through matrix algebra," Computer Physics Communications, vol. 181, pp. 21–23, 2010.
[2] R. J. Onega, "Radioactivity calculations," American Journal of Physics, vol. 37, no. 10, pp. 1019–1022, 1969. [Online]. Available: http://link.aip.org/link/?AJP/37/1019/1

[3] D. S. Pressyanov, "Short solution of the radioactive decay chain equations," American Journal of Physics, vol. 70, pp. 444–445, 2002.

[4] L. Moral and A. F. Pacheco, "Algebraic approach to the radioactive decay equations," American Journal of Physics, vol. 71, pp. 684–686, 2003.

[5] T. M. Semkow, "Comment on "algebraic approach to the radioactive decay equations," by 1. Moral and A.F. Pacheco" American Journal of Physics, vol. 72, pp. 410–411, 2004.

[6] D. Yuan and W. Kernan, "Explicit solutions for exit-only radioactive decay chains," Journal of Applied Physics, vol. 101, no. 9, p. 094907, 2007.

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Back to the new Decay Code: how decay chains and activation are handled

- α_i needs to be know for each
 isotope in each volume
- Code provides a book-keeping class to automatically keep track of these activations resulting from appropriate G4-processes
- After initial activation simulation activities are calculated
- Result is saved in XML file
- Automatic redefinition of geometry including volumes acting as radioactive background sources

Decay matrix

Isotopes in chain after time t

 $\mathbf{N}(t) = e^{\Lambda t} \mathbf{N}(0)$

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Validation of long term activation code with Space Shuttle Data

- Nal-scintillator crystal stored in locker aboard STS-53 mission
- Irradiation and activation by cosmic rays during mission
- After touchdown: mating with PMT and in-situ measurement of gamma-radiation resulting from decaying isotopes (Truscott and Dyer, 1992+94)
- Much simplified Geant4 simulation: cubic homogeneous crystal in hollow aluminium box

Shuttle image: nasa.gov

Validation of long term activation code with Space Shuttle Data

Long Term Activation of the ATHENA WFI (after 10yr)

Final Estimates for the ATHENA Wide Field Imager Background

 ATHENA background is lower or equal to existing missions

 Long term activation plays a negligible role.

ATHENA background goals are feasible

Selected ATHENA background rates

Scenario	10 ⁻⁴ cts keV ⁻¹ cm ⁻² s ⁻¹
with frame-dropping	6.42±0.15
with 200px exklusion	12.43±0.21
with e-field (+200kV/cm)	2.70±2.67
Long time activation	0.21±0.05

ATHENA compared with other missions

Conclusion

The magnitude of the ATHENA Wide Field Imager background has been estimated using a new radioactive decay code for Geant4

At (0.21±0.05) x10⁻⁴cts keV⁻¹ cm⁻² s⁻¹ it is well below the prompt background rate which has been estimated at (2.7–6.42) x10⁻⁴cts keV⁻¹ cm⁻² s⁻¹

The high-Z layer of the graded-Z shielding and the cold finger have been identified as the most activated parts. The graded-Z layers are themselves well shielded from the detector, the emission from the cold finger is detected by pattern rejection