

### Physics methods for the \_\_\_\_\_\_ simulation of photoionisation

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#### **IEEE NSS 2013**

27 October – 2 November 2013

Seoul, Korea

#### 2013 IEEE NSS/MIC/RTSD

"Beyond Imagination of Future Science"

### Rationale

Simulation physics	New physics	New trends			
The simulation of photon physics is well-established.	New theoretical calculations and parameterisations were released recently.	Very low energy(~eV scale) & micro/nano dosimetry ex. FLUKA, Geant4-DNA, MOCA, OREC/ NOREC, PARTRAC, Penelope, PTB-code, Trion etc.			
→ Is there any quantitative validation?	→ Are they accurate than old one?	→ How accurate are models?			
Droject to validate quantitatively a wide cat of simulation					
modeling options against a large collection of experimental data					
State-of-the-art simulation of photon interactions					
Elastic scattering : <b>Published</b>	<b>Photoelectric effect</b>	Compton scattering pair production			

: first results

### **Photoionisation in Monte Carlo codes**

	Cross sections	Angular distribution	
EGS5	PHOTX	Sauter	
EGSnrc	Storm-Israel Fit to XCOM EPDL97 (subshell)	Sauter	
FLUKA	EPDL97	Sauter	
Geant4	Revised Biggs-Lighthill (Henke) EPDL97	Sauter-Gavrila Same direction	
ITS	Scofield 1973	Fischer+Sauter	
MCNP(X)	EPDL89, EDPL97 ENDFB/IV+Storm-Israel		
Penelope	EPDL97	Sauter (K shell)	

### **Photoionisation in Geant4 9.6**



### **Cross section sources**

Year	Compilation	Energy	Ζ	(sub)Shell	Method
1967-1988	<b>Biggs-Lighthill</b>	10 eV – 100 GeV	1-100	-	parameterised
1992	Brennan-Cowan	30 eV – 700 keV	3-92	-	tabulated
2000	Chantler	10 eV – 433 keV	1-92	К	tabulated
2003	Ebel	1 keV – 300 keV	1-92	all	parameterised
2002	Elam	100 eV – 1 MeV	1-98	-	tabulated
1997	EPDL97 (Scofield)	10 eV – 100 GeV	1-100	all	tabulated
1982-1993	Henke	10 eV – 30 keV	1-92	-	tabulated
1970- <mark>2006</mark>	McMaster/Shaltout	1 keV – 700 keV	1-94	-	tabulated
1989	PHOTX (Scofield)	1 keV – 100 MeV	1-100		tabulated
2001	RTAB	10 eV – 30 keV	1-99	all	tabulated
1973	Scofield	1 keV – 1.5 MeV	1-100	all	tabulated
1970	Storm-Israel	1 keV – 100 GeV	1-100	-	tabulated
1973	Veigele	100 eV – 100 MeV	1-94	-	tabulated
1987- <mark>2010</mark>	XCOM (Scofield)	1 keV – 100 GeV	1-100	_	tabulated

Different methods and IEEE NSS 2013 calculations

e.g. Chantler's exchange potential in his DHF calculation is different from Scofield's 5

### Strategy

- Evaluate a large number of available modeling options
- Suitable for use in Monte Carlo simulation codes
  - Tabulated theoretical calculations
  - Simple analytical formulations, with documented parameters
- All options evaluated in the same computational environment
  - Minimize dependencies on other software parts (not always components)
- Quantitative, objective evaluation based on statistical methods
- Establish state-of-the-art for the simulation of photoionisation on objective ground
- Computational performance measured along with physical accuracy

\* We only focus on the simulation with non-polarised photons case.

### **Computational environment**

#### Streamlined software design consistent with Geant4 kernel



### **Experimental data**

#### Collected from the literature

- Total cross sections
- Partial cross sections
- Angular distributions
- ature > 150 references > 5000 data points ~ 3700  $\sigma_{total}$ 
  - ~ 1400  $\sigma_{\text{shell}}$



- Data types
  - Pure experimental cross sections: direct measurements
  - Semi-empirical cross sections: involve theoretical manipulations
    - ▶ e.g. subtraction of calculated scattering contribution (Compton and elastic)

#### Format

- Tables, text
- Figures: digitized, digitization error estimated

#### Evaluation of experimental data

- Systematic effects: identified whenever possible
- Outliers



Number of standard deviations

Number of standard deviations

# Systematic effect?

Difference between calculated and "experimental" total cross sections, expressed in terms of number of standard deviations: **pure experimental** and **semi-empirical** data

Only pure experimental data used in the validation process

Number of standard deviations

Number of standard deviations

### **Data analysis method**

#### Two-stage statistical analysis

- Compatibility of each cross section calculation method with experiment
- 2. Comparison of compatibility with experiment across modeling categories
- Quantitative appraisal of capabilities and differences

Difference across categoriesDifference across categoriesContingency tables $\chi^2$  test<br/> $\alpha = 0.01$ Fisher exact test<br/>Barnard test<br/>Pearson  $\chi^2$  test<br/> $\alpha = 0.05$ 

**efficiency**= **N**↓*pass* /**N**↓*test cases* 

as appropriate

### **Results - Total cross sections**



 Most calculation methods exhibit similar compatibility with experiment for E>250 eV

- Chantler, Brennan-Cowan look worse
- Degraded accuracy below 250 eV

	Analysis of contingency tables			
		EPDL Chantler	EPDL Brennan-Cowan	
	Fisher	0.044	0.011	
	Pearson $\chi^2$	0.033	0.007	
IEEE NSS	Barnard	0.035	0.007	



### **Results - Shell cross sections**





Systematic effect observed with RTAB shell cross sections (presumably a missing factor in the calculation) IEEE NSS 2013

Calculated **inner shell** cross sections compatible with experiment

**Outer shell** cross sections inconsistent with experimental data *Beware: small data sample, limited data sources* 

#### p-value $\chi^2$ test

shell	EPDL	Chantler	RTAB	scRTAB	Ebel
K	0.209	0.350	< 0.001	0.315	<0.001
L1	0.075		< 0.001	0.069	0.964
L2	0.339		< 0.001	0.299	0.154
L3	1		< 0.001	1	1
M1	<0.001		< 0.001	<0.001	
M4	0.031		< 0.001	<0.001	
M5	<0.001		< 0.001	<0.001	
N1	<0.001		< 0.001	<0.001	
N6	<0.001		< 0.001	<0.001	<0.001
N7	<0.001		< 0.001	<0.001	<0.001
01	<0.001		< 0.001	<0.001	<0.001
02	<0.001		< 0.001	<0.001	<0.001
O3	<0.001		<0.001	<0.001	<0.001
P1	<0.001		<0.001	<0.001	<0.001

### **Results - Angular distribution**



### Conclusion

- Large scale effort to evaluate quantitatively physics methods for photoionisation simulation
  - Part of a wider project for **quantitative** assessment of **state-of-the-art** simulation of photon interactions

#### Total cross section

- Most calculation methods exhibit similar behaviour
- More recent calculations (Chantler, Brennan-Cowan) do not appear more accurate than old Scofield's 1973 (unrenormalized)

#### Inner shells

- EPDL, *(corrected)* RTAB appear equivalent, Ebel's parameterisation inconsistent with experimental K shell data

#### Outer shells

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- No calculation method appears adequate to reproduce experimental data

#### Photoelectron angular distribution

- Scarce data and experimental systematics prevent a quantitative discrimination

#### All results will be documented in detail in a forthcoming publication















## ...a big **THANK YOU** to the CERN Library!